

Dark Matter in Pop III stars: The effect of a hypothetical object over another

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Introduction

Population III stars are the **first supplier of light and metal of the universe**. As formed in dense **Dark Matter (DM) halos**, their formation and evolution can be strongly affected by the **capture and annihilation of Weakly Interacting Massive Particles (WIMPs)** a DM candidate.

In this work we run models of $20 M_{\odot}$ with **different WIMPs densities and rotation velocities**, thanks to the stellar evolutionary code GENE (observatoire de Genève), and analyse their effects on the structure, lifetime and chemical composition of the star along its evolution (Pauchet et al. 2026).

Baryon-DM interaction

DM can interact with baryonic matter through two main mechanisms:

Capture: a WIMP passing by the star will be scattered by baryons until it loses enough momentum and is captured by one of them.

Annihilation: as WIMPs are their own antiparticle they can annihilate themselves which bring energy to the star.

→ take place in the **densest part i.e. the stellar core**

WIMPs annihilation produces a **new source of energy throttling nuclear burning** ⇒ the star is sustained without using its Hydrogen fuel ⇒ **extended Main Sequence (MS)** (see Figure 1)

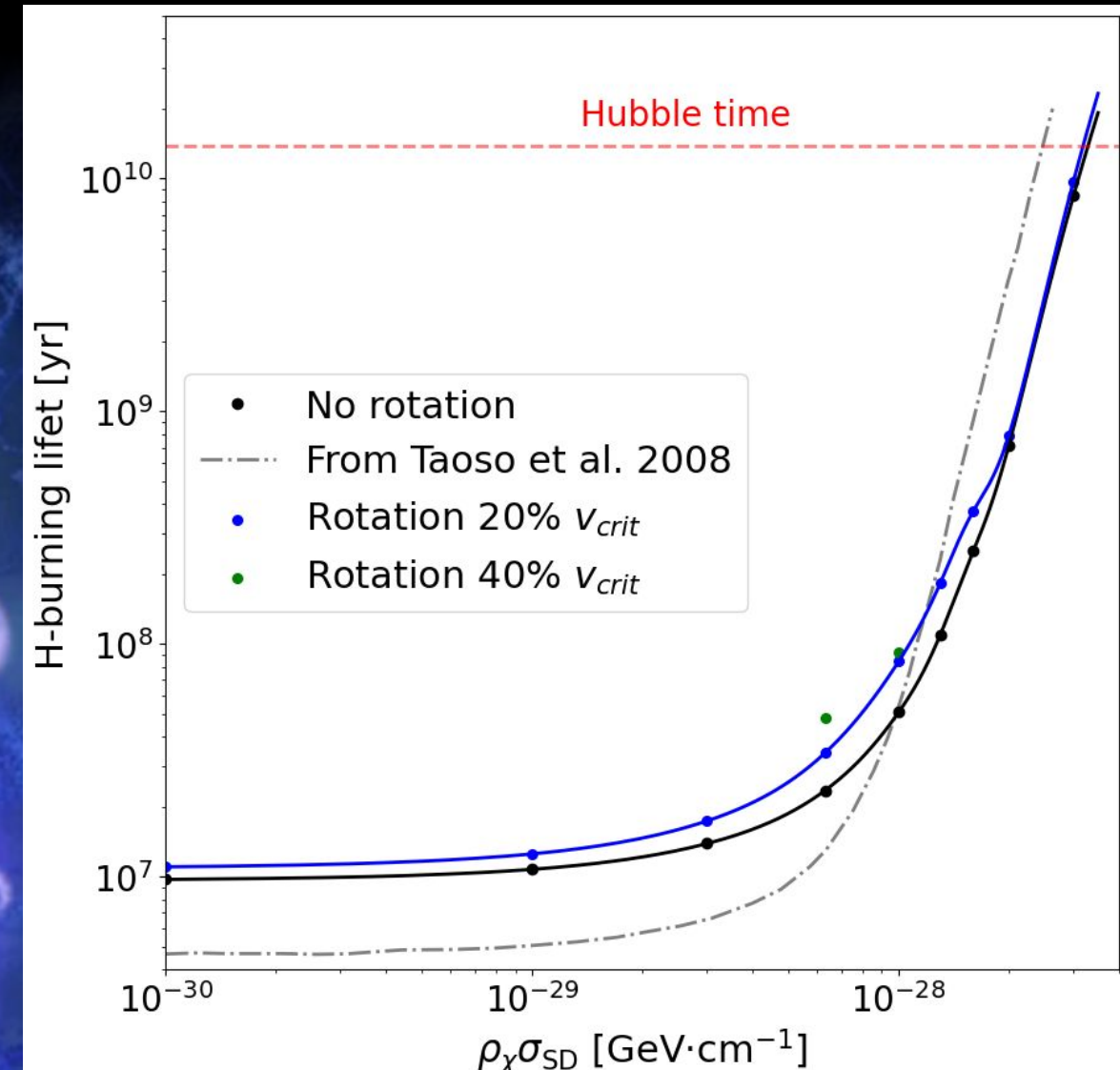
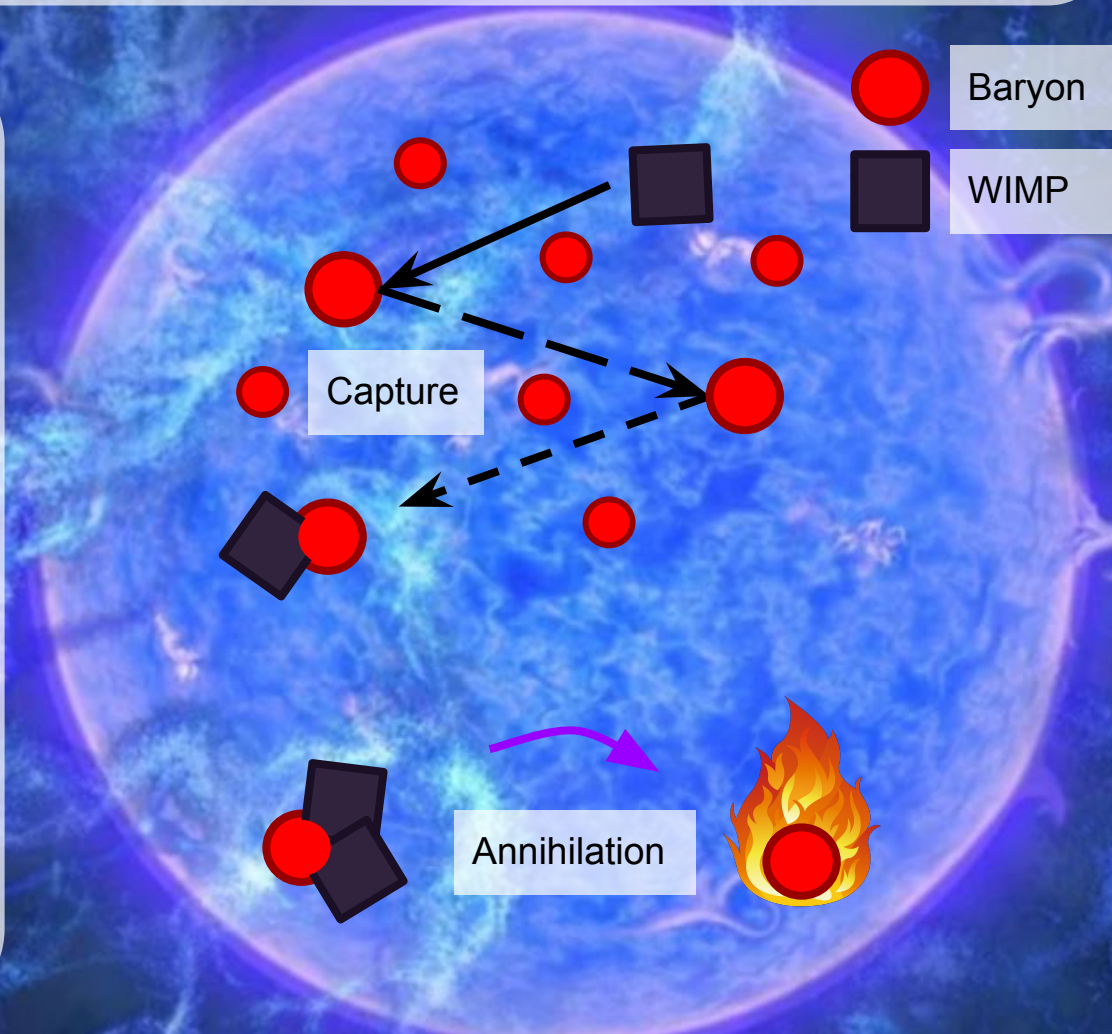


Figure 1: MS duration for a $20 M_{\odot}$ stars as a function of the WIMPs density surrounding the star ($\sigma_{SD} = 10^{-40} \text{ cm}^2$). Blue and green dots describe rotating stars at 20% and 40% of the critical velocity (v_{crit}).

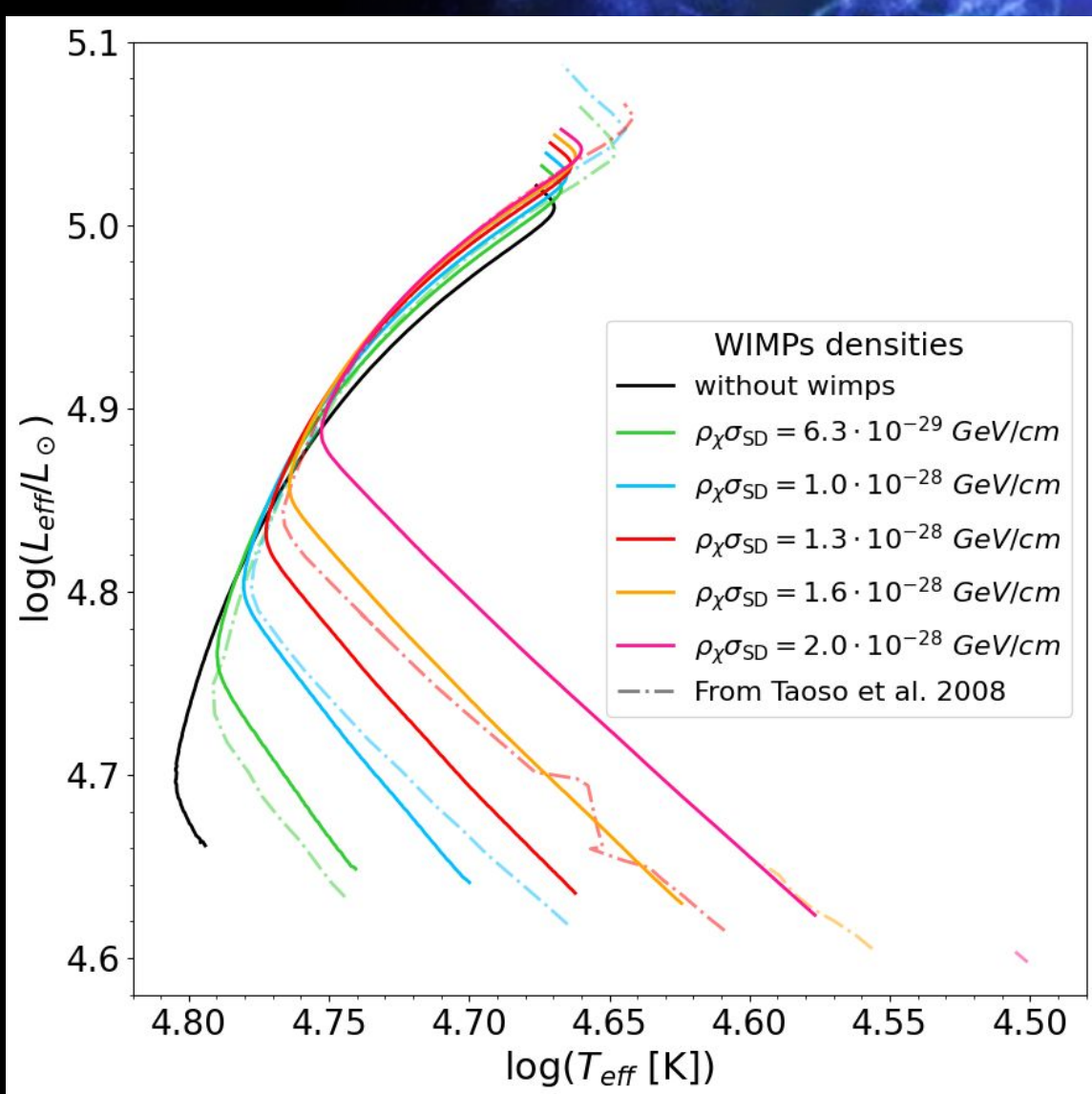


Figure 2: Hertzsprung-Russel Diagram for a $20 M_{\odot}$ stars within a halo of different WIMPs densities

Effect on the star structure

At the ZAMS: additional energy injection from WIMPs annihilation in the stellar core **inflates the star** → **larger radius (cooler Teff)**.

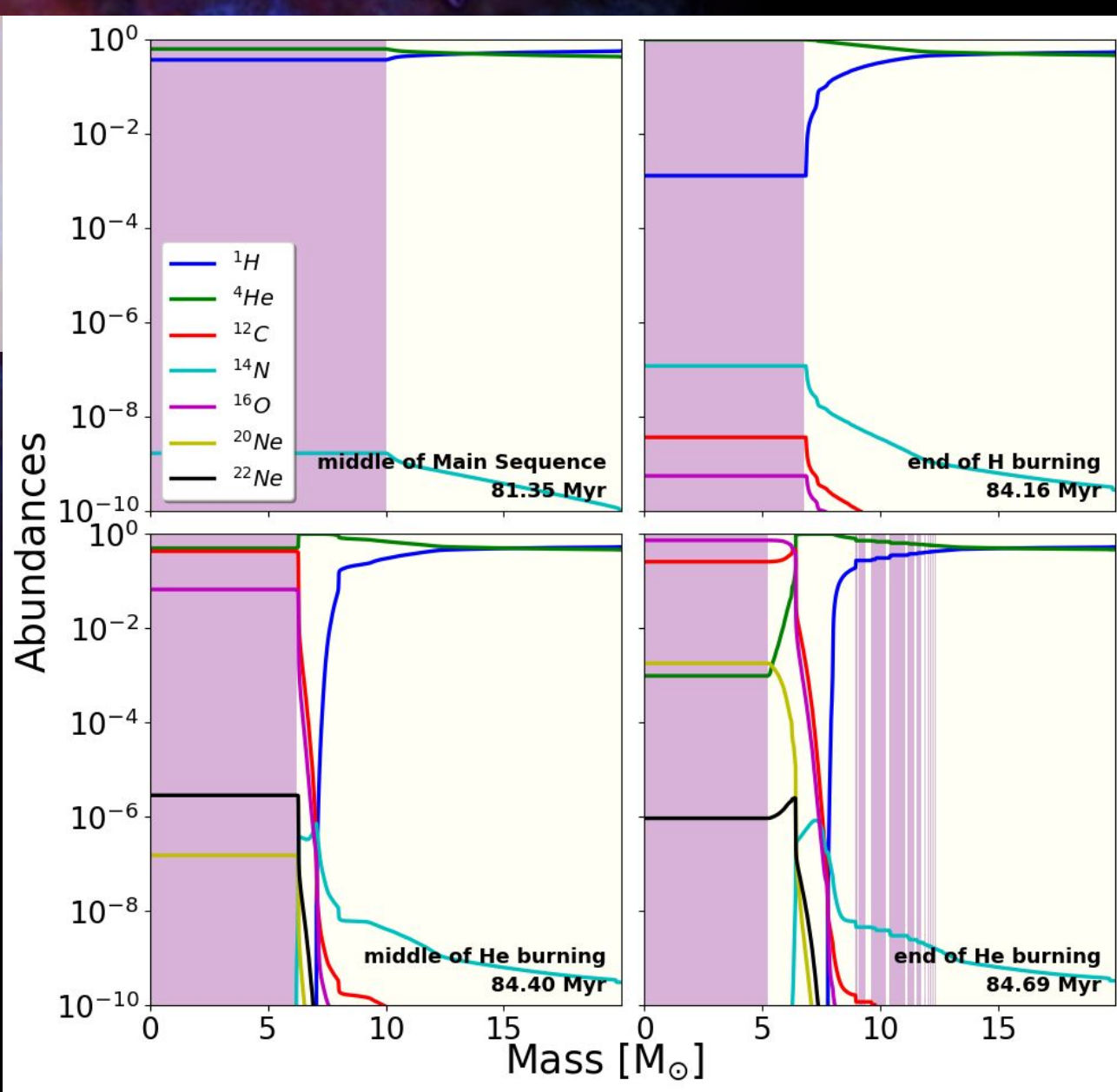
⇒ **energy output is dominated by WIMP annihilation rather than nuclear fusion**, the star undergoes only minimal H-burning during MS.

As evolution proceeds, the **DM energy input gradually becomes less dominant**. Once a significant amount of H is converted into He in the core, the **WIMP capture rate drops** → nuclear energy production steadily increases and **the star resumes a more conventional MS evolution** (see Figure 2).

Rotating stars

Adding rotation to the system, **even a relative low rotation velocity i.e. 20% of v_{crit}** is enough to show an **extended mixing of chemicals and heat transport towards the surface**. This leads to hotter and more luminous stars even when H-burning becomes dominant again.

Figure 3: Chemical abundances as a function of the mass enclosed inside a specific radius of a $20 M_{\odot}$ star at 20% of v_{crit} with DM. Each plot shows a different star evolution stage. The pink area describes the helium core area.



Chemical mixing and SuperNova (SN) yields

Surface abundances: For a DM-star with $0.2 v_{crit}$, H surface abundance is lowered to ~ 0.53 while He increases to $Y \sim 0.47$ in comparison to $X \sim 0.75$ and $Y \sim 0.25$ for a DM-free star, **values generally reached for DM-free stars at much larger spin $> 0.5 v_{crit}$** . The same model also leads to **extreme CNO yields, i.e. 50 times higher surface abundances**.

SN yields: The integrated masses, from the He core to the surface **are 45% poorer in H and 33% richer in He for the DM + rotation model**. Additionally the same model gives SN yields **5 and 8 times more massive for N and O, respectively**.

Conclusion

- **lifetime:** extension of MS which can reach billions of years.
 - **stellar structure:** At the ZAMS the static WIMP model is 26% cooler and 55% larger compared to the DM-free model.
 - **mixing:** rotation + extended lifetime → surface abundances richer in metals and poorer in H.
 - **SN yields:** rotating model with WIMPs = most promising model as it exhibits an \nearrow in He, N and O, while \searrow for H.
- Stars formed in a DM halo could be **the progenitors of massive black holes in the center of galaxies**. The challenge is **how to detect them?** With this work we hope to open the research on those fascinating objects.