

Wolf-Rayet evolution computations with realistic atmospheres

**VOJE, T. and
PALACIOS, A. and MARTINS, F.**
LUPM, Univ. Montpellier, CNRS,
Montpellier, France

Mail : thomas.voje@umontpellier.fr

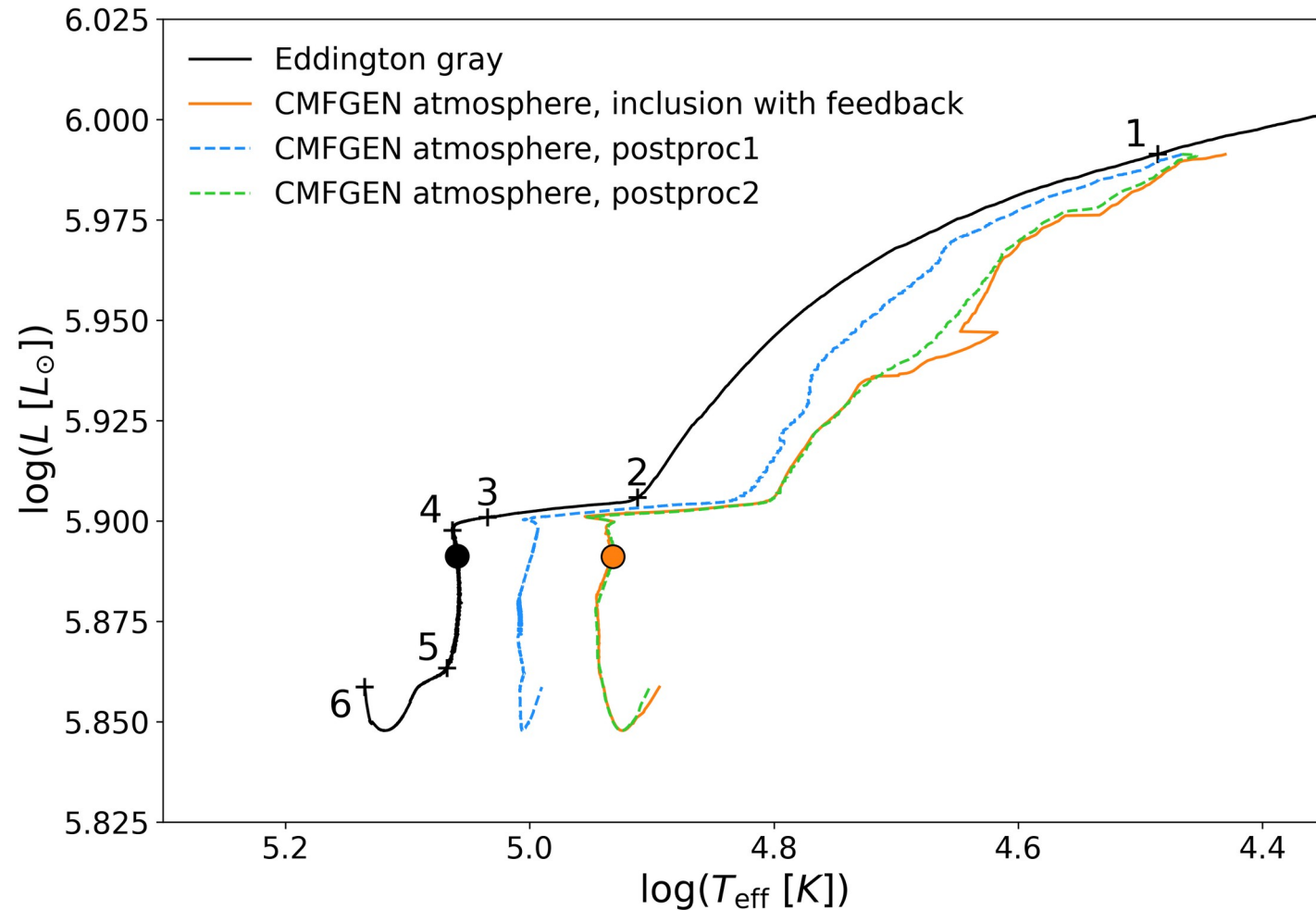


Figure : Evolutionary track of a $60 M_{\odot}$ stellar evolution model at solar metallicity with a CMFGEN atmosphere (orange) compared to an identical model with an Eddington gray atmosphere (black). Models computed using post-processing methods are also shown for comparison (blue and green).

